Improvement of DWT-SVD with Curve Fitting and Robust Regression: An Application to Astronomy Images

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Abstract—DWT-SVD is a frequency domain based eigenanalysis watermarking technique. In this work, we improve this method by exploring the relationship between the cover image’s DWT singular values and those of the watermark. We show that, via the usage of curve fitting and robust regression, it is possible to achieve accurate results. We also demonstrate that the improved scheme is suitable for the watermarking of astronomy images. In addition to encoding and decoding examples, statistical results on stealth and robustness are deduced from the experiments so that the clear advance can be observed. Quality of the watermark is measured by testing against various attack types.

Index Terms—Watermarking; discrete wavelet transforms; eigenvalues and eigenfunctions; curve fitting.

I. INTRODUCTION

Digital watermarking has a wide area of application. The most important one of these is copyright security; a secret message is embedded into the media so that the owner knows it stands as a signature. For the battle against piracy and for the management of media tracking, it is impossible to neglect the benefit of information hiding [1]–[5].

The domain of visual watermarking is classified into two types. The first one is the obvious one where the pixels reside; spatial domain. The second one is the frequency domain, where we embed the secret message into the frequency coefficients obtained via an analysis of the cover image. Discrete Wavelet Transform (DWT) and Discrete Cosine Transform (DCT) are examples of the state-of-the-art frequency domain analysis. In basic DWT embedding, the visual watermark is hidden into the wavelet coefficients of the cover image [6]–[10]. That is, coefficients are modified by using the intensity values of the watermark. In DWT-SVD embedding, singular values of the frequency coefficients are modified by using the singular values of the visual watermark. In this work, we focus on the improvement of DWT-SVD technique [11].

Although the central technique DWT is somehow old, its modified and upgraded versions, that is, strengthened ways of frequency domain analysis via SVD or other transformations, are still in the field of active research [12], [13]. Moreover, even DWT itself is engineered in more articulated techniques [14], which is a proof of the solid state of this classical algorithm. Variants such as DCT are also up-to-date base selections for implementations of watermarking [15].

Influenced by [10], we have derived a technique which will be explained as follows: in II.A base theoretical tools, in II.B and II.C extensions such as curve fitting are noted. Overall scheme development and experiment details are given in II.D, IIE, respectively.

II. IMPROVEMENT OF DWT-SVD

Our aim at this work was to establish an improvement of DWT-SVD [10], [11], [14] via an additional analysis on singular values. For this, we utilized curve fitting and robust regression. Thus, the derived technique can be summarized as DWT-SVD-CF-RR (DSCR). The ideal advance of a watermarking scheme must be in both robustness and stealth. Robustness is the ‘strength’ measure of the encoding to a range of attacks. Stealth, on the other hand, is the transparency of the hidden message; this is essential for the overall commercial quality of the modified image.

Before carrying out the experiments, our main idea was this: what if one embeds an approximation of watermark singular values as a function of the cover DWT singular values rather than directly using the initial (original) ones? This corresponds to a more consistent modification since the same function is used for the band’s all singular values; the resulting sequence is retrieved by a well-defined single variable function. On the other hand, since an approximation is used, the connection to the watermark is almost retained. Hence, our initial guess was that the final PSNR value of the encoded image should be higher than that of the standard DWT-SVD’s output and the correlation values of the decoded singular values should have a tolerable distance to those of the standard strategy. After experiments, we have seen that, both PSNR values and correlation measurements were better.

A. DWT-SVD

Assume that we have a $2n \times 2n$ cover and an $n \times n$
watermark image.

1) Embedding

1. Apply DWT and get subbands LL, LH, HL and HH.
2. Apply SVD to the watermark image and get
\[ U_w \Sigma_w V_w^T \] \[ \Sigma_w \] is a diagonal matrix containing the
singular values \( \lambda_{w1}, \lambda_{w2}, \ldots, \lambda_{wn} \). \( U_w \) and \( V_w \) are
orthogonal matrices.
3. For i-th sub band apply SVD and get the
decomposition \( U_i \Sigma_i V_i^T \). Make this for all i = 1, 2, 3, 4,
where i = 1, 2, 3, 4 corresponds to indices of LL, LH, HL and HH respectively.
4. Modify the subbands as \( U_i (\Sigma_i + q_i \Sigma_w) V_i^T \). \( q_i \) is a
scaling factor. That is, modify the singular values of i-th sub band as
\( \lambda_{i1} + c_i \lambda_{w1}, \lambda_{i2} + c_i \lambda_{w2}, \ldots, \lambda_{in} + c_i \lambda_{wn} \)
and apply inverse SVD transform.
5. Apply inverse transform to the subbands to get the
watermarked image.

2) Extraction

1. Apply DWT to the input image \( I^* \) and get subbands LL, LH, HL and HH.
2. For i-th subband apply SVD and get the decomposition
\[ U^*_i \Sigma^*_i V^*_i \] \[ V^*_i \] is a diagonal matrix containing the values
\( \lambda^*_i \). Make this for all i = 1, 2, 3, 4, where i = 1, 2, 3, 4 corresponds to LL, LH, HL and HH respectively.
3. Find the constructed singular values \( \left( \lambda^*_i - \lambda_{wi} \right) / \alpha_i \) for
j = 1, 2, ..., n.
4. Output the visual watermark as \( U_w \Sigma^*_w V_w^T \) where \( \Sigma^*_w \) is a
diagonal matrix containing the values \( \left( \lambda^*_i - \lambda_{wi} \right) / \alpha_i \) for j=1, 2, ..., n.

B. Curve Fitting

Assume that we have m points \( \{(x_1,y_1),(x_2,y_2),\ldots,(x_m,y_m)\} \). We do at (least-squares) curve fitting is to find a d degree polynomial
\[ f_i(x) = a_dx^d + a_{d-1}x^{d-1} + \ldots + a_1x + a_0, \] (1)
such that \( \sum_{i=1}^{m} (f_i(x_i) - y_i)^2 \) is minimized. We formulated
curve fitting for the LL band as follows: for a given degree d, find \( f_i(x) \) such that \( \sum_{i=1}^{n} (f_i(x_i) - y_i)^2 \) is
minimized. Hence, we fit a polynomial of LL band singular values to approximate the singular values of the watermark.

C. Robust Linear Regression

Ordinary Least Squares (OLS) is formulated as follows: given \( X, y \) matrices where the dimensions are \( n \times D \) and
\( n \times 1 \) respectively, find \( D \times 1 \) dimensional matrix \( \beta = \hat{\beta} \) such that
\[ X\beta - y, \] (2)
is minimized. That is, we model the relationship between an
input vector \( x \) and a target value \( y \) in the form of a linear function characterized by \( \beta \), where the prediction \( \hat{y} \) is
evaluated as
\[ \hat{y} = x_1 \beta_1 + x_2 \beta_2 + \ldots + x_D \beta_D, \] (3)
so that the total distance of the estimated values to the actual ones is minimized.

While modeling via OLS, we do not take the weights of the samples into account. If we do it, the result is a new formulation which contains the analysis of outliers and leverage points; Weighted Least Squares (WLS). Outliers are points that are not fitted very well by the linear model (i.e. points having large residuals), whereas leverage points are those outlying in the x-space [3]. Since the weight assigning procedure is dependent to residuals and residuals are calculated at each iteration, weighting is done iteratively. Robust Regression can be summarized as solving
\[ \arg\min_{\beta} \sum_{i=1}^{n} |y_i - x_i^T \beta|^2 \] (4)
at each j-th iteration. Weighting is done according to the residuals of the last fit
\[ r^{j-1} = y - y^{j-1} = y - X\hat{\beta}^{j-1}, \] (5)
and the leverage matrix
\[ H = X(X^TX)^{-1}X^T. \] (6)

D. Scheme

Assume that we have a \( 2n \times 2n \) cover and an \( n \times n \) watermark image.

1) Embedding

Given \( \{\alpha_1, \alpha_2, \alpha_3, \alpha_4, t\} \in \mathbb{R}^+, s = [s_1s_2s_3s_4]^t \), \( d \in \mathbb{N} \)
where \( \alpha_i \) are the scaling factors of DWT- SVD, \( t \) is the
tuning constant of robust regression, \( s \) is a binary string (i.e. \( s_i \in \{0,1\} \)) and \( d > 1 \) is the degree of the curve fitting polynomial.
1. Apply DWT and get subbands LL, LH, HL and HH.
2. Apply SVD to the watermark image and get
\[ U_w \Sigma_w V_w^T \] \[ \Sigma_w \] is a diagonal matrix containing the
singular values \( \lambda_{w1}, \lambda_{w2}, \ldots, \lambda_{wn} \). \( U_w \) and \( V_w \) are
orthogonal matrices.
3. For i-th subband apply SVD and get the decomposition
\( U_i \Sigma_i V_i^T \). Make this for all i = 1, 2, 3, 4, where i = 1, 2, 3, 4 corresponds to indices of LL, LH, HL and HH respectively.
4. Find the polynomial fitting function \( f_i \) of degree d, where
the input set is \( \{\lambda_{i1}, \lambda_{i2}, \lambda_{i3}, \lambda_{i4}\} \), i.e. fit a function of singular values of the LL band with the output values of the watermark’s singular values.
5. For \( i \in \{2,3,4\} \) if \( s_i = 1 \), use robust regression to find
a least-squares approximation \( f_i \) on
Histogram Equalization

A greater PSNR of the encoded image and a higher mean correlation value – couldn’t be found, we modified $s$ gradually; we first tried $s=1110$, later $s=1101$ and so on. Here, $s$ is of length 4 for the sake of convenience; we never omitted curve fitting – approximation embedded in LL band – modification. Choosing the high-PSNR values of the original DWT-SVD algorithm, we set $\alpha_1=0.05$, $\alpha_2=\alpha_3=\alpha_4=0.005$. We used bisquare weighting and the same $t$ for all robust regression steps.

Reported mean correlation and PSNR values are obtained when DSCR got a better result compared to DWT-SVD, i.e. these are not the best values, albeit the superior outcome.

2) Astronomy Images

This work started with an application-driven idea of implementing a visual watermark algorithm to the domain of astrophotography. We used Hubble Site [4] for the dataset of images.

Examples of deep sky images are shown in Fig. 1.

- **Messier 101**: a spiral galaxy (Fig. 1(a)). Estimated number of stars it consists is about one trillion.
- **NGC 290**: a star cluster in the Small Magellanic Cloud (Fig. 1(b)).
- **Messier 74**: also known as NGC 628, a spiral galaxy slightly smaller than Milky Way (Fig. 1(c)).
- **LH 95**: a star forming region of glowing hydrogen in the Large Magellanic Cloud [3] (Fig. 1(d)).

Using these images and performing a sequence of 10 attacks: JPEG 75 (JPEG compression with quality 75), JPEG 50 (JPEG compression with quality 50), JPEG 25 (JPEG compression with quality 25), Gaussian Noise (0 mean and 0.001 variance), Mean Filter (2-D filter mean filter), Resize (512x512 → 256x256 → 512x512 bilinear reszing), Rotation (20 degrees), Histogram Equalization (Contrast enhancement via histogram equalization), Intensity Adjustment (Intensity interval [0, 0.8] mapped to [0, 1]), Gamma Correction (Intensity adjustment with gamma = 1).

On the encoded image, we measured PSNR and mean correlation. Each correlation value is recorded as in [9], i.e. by taking the highest of all bands (including the fusion result 2.4.1). Differing from [1], we used mean correlation as a final robustness metric

$$\frac{1}{N} \sum_{i=1}^{N} \text{corr}(\lambda_i, \lambda_w).$$

where $\lambda_i$ and $\lambda_w$ are the i-th reconstructed singular value vector (from the i-th attacked image) and the visual watermark’s singular value vector, respectively. corr is Pearson correlation function.

For NGC 290, the original, encoded and decoded images can be seen in Fig. 2. Detailed results for M101, NGC 290, M74 and LH 95 can be seen in Table I.
most striking result is obtained on NGC 290, where we have an obvious advance especially in terms of robustness: a boosted mean correlation value of 0.8511.

3) Classical Images

From the following results, one can easily see that, the proposed technique is not a specialized-for-deep-sky strategy. In this section we demonstrate DSCR embedding on well-known image processing data (https://homepages.cae.wisc.edu/~ece533/images/).

For Goldhill, the original, encoded and decoded images can be seen in Fig. 4. Detailed results for Airplane, Baboon, Boat, Goldhill, Barbara, Lenna and Watch are shown in the Table II.

### III. Conclusions

In this work, an idea based on expressing the set of watermark singular values by functions of the cover image’s DWT domain singular values is implemented. Since such an approximation enforces a more consistent transform on the cover image’s DWT singular values, a higher PSNR value is obtained. Moreover, together with the utilization of a two-level robust regression, a boosting of the mean correlation value is achieved. Obviously, one can find more sophisticated ways of fusing the components, e.g. through Support Vector Regression, albeit the thematic concentration on iterative reweighted least-squares of this

### TABLE I. DEEP SKY RESULTS

<table>
<thead>
<tr>
<th>Image</th>
<th>DSCR PSNR</th>
<th>DSCR MC</th>
<th>d</th>
<th>t</th>
<th>s</th>
<th>DS PSNR</th>
<th>DS MC</th>
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<tbody>
<tr>
<td>M101</td>
<td>37.5286</td>
<td>0.9614</td>
<td>9</td>
<td>0.1</td>
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<td>1100</td>
<td>37.4955</td>
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<td>2.0</td>
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<td>11</td>
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<td>M74</td>
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<td>0.7</td>
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<td>110</td>
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<td>LH 95</td>
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<td>8</td>
<td>0.1</td>
<td>11</td>
<td>111</td>
<td>37.4955</td>
</tr>
</tbody>
</table>

Note: DS: Original DWT-SVD; DSCR: DWT-SVD with Curve-Fitting and Robust Regression; MC: Mean correlation; d: Degree of the DSCR polynomial-fitting; t: Tuning constant for DSCR robust regression; s: Binary string of DSCR (2.D.1).

Fig. 1. Deep sky images [4]: a) M101; b) NGC 290; c) M74; d) LH 95.

Fig. 2. Embedding & extraction for NGC 290: (a) NGC 290; (b) encoded (PSNR = 37.6242); (c) watermark; (d) decoded.

Fig. 3. Embedding & extraction for Goldhill: (a) goldhill; (b) encoded (PSNR = 37.6144); (c) watermark; (d) decoded.

Fig. 4. Embedding & extraction for Goldhill: (a) goldhill; (b) encoded; (c) watermark; (d) decoded.

### TABLE II. RESULTS.

<table>
<thead>
<tr>
<th>Image</th>
<th>DSCR PSNR</th>
<th>DSCR MC</th>
<th>d</th>
<th>t</th>
<th>s</th>
<th>DS PSNR</th>
<th>DS MC</th>
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<td>0.8035</td>
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<td>Boat</td>
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<td>Goldhill</td>
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<td>0.8932</td>
<td>8</td>
<td>0.1</td>
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<td>Barbara</td>
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<td>Lenna</td>
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<td>Watch</td>
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</table>

- Airplane
- Baboon
- Boat
- Goldhill
- Barbara
- Lenna
- Watch

For Goldhill, the original, encoded and decoded images can be seen in Fig. 4. Detailed results for Airplane, Baboon, Boat, Goldhill, Barbara, Lenna and Watch are shown in the Table II.

Hence, we have an improvement of DWT-SVD via curve-fitting and robust regression for each astronomy image. The
scheme.

REFERENCES


